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We have to analyze several CCD images from a star field to perform photometry on an exoplanet transit over its star.

*(Observation of HD189733/b on July, 18<sup>th</sup>, 2017 at the 80 cm telescope of "Osservatorio Polifunzionale del Chianti"<sup>1</sup>)*

- Observation data:

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- Exoplanet detection is based on a decrease of measured flux of the star due to the occultation of the planet passing in front of the star.

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- We do not need absolute (calibrated) flux values
- In order to remove the effects of flux variations due to other causes (e.g.: variation of air transparency) we compute the flux of the target star relative to the flux of other stars in the same image.
- In other terms, given the flux of the target star  $F_T$  and the fluxes of some other stars from the same image,  $F_1, F_2, \dots, F_N$ , the relative flux is computed as:

$$F_R = \frac{F_T}{\sum_{i=1}^N F_i}$$

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# Data Analysis Pipeline -1

A complete example - 3

## class Pipeline

All the procedures will require to operate on the full sequence of images.

The class **Pipeline** has the purpose to apply a specified function to a selection of (possibly all) the data files in a given directory.

### file: pipeline.py

```
import os, json, glob, sys, getopt
from astropy.io import fits
import numpy as np

class Pipeline:
    def __init__(self, operate, common_args=None):
        ddir, glbn, first, number, step = getopt.getopt(sys.argv[1:], "d:g:f:n:s:")
        self.filenames = glob.glob(os.path.join(ddir, glbn))
        self.filenames.sort()
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get\_args(): see below

See module: glob

Manage added arguments

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Annotations:

- get\_args(): see below**: Points to the `get_args()` call in the `__init__` method.
- See module: glob**: Points to the `glob` import statement.
- Manage added arguments**: Points to the `common_args` parameter in the `__init__` method.
- Manage options -f (first) e -n (number)**: Points to the `first` and `number` parameters in the `__init__` method.

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Annotations:

- `get_args(): see below` (points to `get_args()` call)
- `See module: glob` (points to `glob.glob`)
- `Manage added arguments` (points to `common_args` assignment)
- `Manage options -f (first) e -n (number)` (points to `first` and `number` filters)
- `Function to apply to each data file` (points to `self.operate` assignment)

# Data Analysis Pipeline -1

A complete example - 3

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    def _getdata(self, filename):
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Annotations:

- `get_args(): see below`
- `See module: glob`
- `Manage added arguments`
- `Manage options -f (first) e -n (number)`
- `Function to apply to each`
- `Method _getdata gives an error if not reimplemented`

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Annotations:

- `get_args(): see below`
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- `Manage options -f (first) e -n (number)`
- `Function to apply to each`
- `Method _getdata gives an error if not reimplemented`
- `Loop on the data file sequence`

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Annotations from top to bottom:

- get\_args(): see below
- See module: glob
- Manage added arguments
- Manage options -f (first) e -n (number)
- Function to apply to each
- Method \_getdata gives an error if not reimplemented
- Loop on the data file sequence
- The function is called here

# Data Analysis Pipeline -1

A complete example - 3

## class Pipeline

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Manage added arguments

get\_args(): see below

See module: glob

Manage options -f (first) e -n (number)

Function to apply to each

Method `_getdata` gives an error if not reimplemented

Loop on the data file sequence

The function is called here

Manage option -s (step)

# Data Analysis Pipeline -2

A complete example - 4

## function get\_args()

Has the purpose to manage command line arguments for programs using module pipeline.

file: pipeline.py

```
def get_args():
    try:
        opts, args = getopt.getopt(sys.argv[1:], 'd:g:f:n:sh')
    except getopt.GetoptError:
        print("Argument error (-h: help)")
        sys.exit()

    ddir = "."
    glbn = "*.*"
    first = None
    number = None
    step = False
    for (o, v) in opts:
        if o == '-h':
            print(__doc__)
            sys.exit()
        elif o == '-d':
            ddir = v
        elif o == "-g":
            glbn = v
        elif o == "-f":
            first = int(v)
        elif o == "-n":
            number = int(v)
        elif o == "-s":
            step = True
    return (ddir, glbn, first, number, step)
```

# Data Analysis Pipeline -2

A complete example - 4

## function get\_args()

Has the purpose to manage command line arguments for programs using module pipeline.

file: pipeline.py

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def get_args():
    try:
        opts, args = getopt.getopt(sys.argv[1:], 'd:g:f:n:sh')
    except getopt.GetoptError:
        print("Argument error (-h: help)")
        sys.exit()

    ddir = "."
    glbn = "*.*"
    first = None
    number = None
    step = False
    for (o, v) in opts:
        if o == '-h':
            print(__doc__)
            sys.exit()
        elif o == '-d':
            ddir = v
        elif o == "-g":
            glbn = v
        elif o == "-f":
            first = int(v)
        elif o == "-n":
            number = int(v)
        elif o == "-s":
            step = True
    return (ddir, glbn, first, number, step)
```

# Data Analysis Pipeline -2

A complete example - 4

## function get\_args()

Has the purpose to manage command line arguments for programs using module pipeline.

file: pipeline.py

```
def get_args():
    try:
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See module **getopt**

# Data Analysis Pipeline -2

A complete example - 4

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file: pipeline.py

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See module **getopt**

Default values

# Data Analysis Pipeline -2

A complete example - 4

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See module **getopt**

Default values

Get optional arguments

# Data Analysis Pipeline -2

A complete example - 4

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Return arguments

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A complete example - 4

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    return (ddir, glbn, first, number, step)
```

See module **getopt**

Default values

Get optional arguments

Return arguments

The above function allows all programs using function get\_args() to accept the same command line arguments in a consistent way.

### Deriving useful classes from Pipeline

Class Pipeline is only a framework: in order to do something useful its `_getdata()` method must be implemented.

Actually, the base class method, if used as it is, generates an error: this is wanted, so that the user is reminded that he must derive from Pipeline and provide a proper implementation.

#### file: pipeline.py

```
class RawPipeline(Pipeline):
    "Read raw data files"
    def _getdata(self, filename):
        data, hdr = fits.getdata(filename, 0, header=True)
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#### file: pipeline.py

```
class CalPipeline(Pipeline):
    "Read calibration files"
    def _getdata(self, filename):
        data = utils.load(filename)
        return data, {"filename": filename}
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#### file: pipeline.py

```
class CalPipeline(Pipeline):
    "Read calibration files"
    def _getdata(self, filename):
        data = utils.load(filename)
        return data, {"filename": filename}
```

In both the above classes the `_getdata()` method is called with the filename as argument and returns data from the file and some auxiliary information, e.g.: the file name.

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        return data, {"hdr": hdr, "filename": filename}
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class CalPipeline(Pipeline):
    "Read calibration files"
    def _getdata(self, filename):
        data = utils.load(filename)
        return data, {"filename": filename}
```

In both the above classes the `_getdata()` method is called with the filename as argument and returns data from the file and some auxiliary information, e.g.: the file name.

### Properly showing the image

The following utility function will be used several times in our procedures

#### file: show.py

---

```
import matplotlib.pyplot as plt
from astropy.visualization import SqrtStretch
from astropy.visualization.mpl_normalize import ImageNormalize

def show(img):
    norm = ImageNormalize(stretch=SqrtStretch())
    plt.imshow(img, cmap="tab10", norm=norm)
```

---

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The module `astropy.visualization` provides useful tools for proper visualization of astronomical images

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The module `astropy.visualization` provides useful tools for proper visualization of astronomical images

## Using `show()`

```
In [1]: from show import show
In [2]: from astropy.io import fits
In [3]: img = fits.getdata("img-459.fit")
In [4]: show(img)
```

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```
import matplotlib.pyplot as plt
from astropy.visualization import SqrtStretch ←
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### Using `show()`

```
In [1]: from show import show
In [2]: from astropy.io import fits
In [3]: img = fits.getdata("img-459.fit")
In [4]: show(img)
```



# Data calibration -1

A complete example - 7

To perform data calibration we use files: `masterdark.fits` and `masterflat.fits`. The files are already preprocessed so that the calibrated image ( $C$ ) is generated as:

$$C = \frac{(R - D)}{F}$$

where  $R$  is the raw image  $D$  is the *dark* and  $F$  is the *flat*.

file: `calibrate.py`

```
import os
from pipeline import RawPipeline, fits
from numpy import ma

DATADIR = "/dati-18.7.2017"

MIN_DARK = -15.0
MAX_DARK = 15.0
MIN_FLAT = 0.85
MAX_FLAT = 1.10

def calibrate(img, step, args, common_args):
    shortname = os.path.basename(args["filename"])
    print("--- Calibrating:", shortname, end=" ", flush=True)
    cal = (img-common_args["dark"])/common_args["flat"]
    calname = os.path.join("work", shortname.replace("fit", "cal"))
    cal.dump(calname)
    print("OK")

if __name__ == "__main__":
    dfile = os.path.join(DATADIR, "masterdark.fits")
    ffile = os.path.join(DATADIR, "masterflat.fits")
    dark = ma.masked_outside(fits.getdata(dfile), MIN_DARK, MAX_DARK)
    flat = ma.masked_outside(fits.getdata(ffile), MIN_FLAT, MAX_FLAT)
    pipe = RawPipeline(calibrate, {"dark":dark, "flat":flat})
    pipe.run()
```

# Data calibration -2

## A complete example - 8

Let's start a description of the procedure from the bottom of file calibrate.py

```
if __name__ == "__main__":
    dfile = os.path.join(DATADIR, "masterdark.fits")
    ffile = os.path.join(DATADIR, "masterflat.fits")

    dark = ma.masked_outside(fits.getdata(dfile), MIN_DARK, MAX_DARK)
    flat = ma.masked_outside(fits.getdata(ffile), MIN_FLAT, MAX_FLAT)

    pipe = RawPipeline(calibrate, {"dark":dark, "flat":flat})

    pipe.run()
```

# Data calibration -2

A complete example - 8

Let's start a description of the procedure from the bottom of file calibrate.py

```
if __name__ == "__main__":
     Direct execution
    dfile = os.path.join(DATADIR, "masterdark.fits")
    ffile = os.path.join(DATADIR, "masterflat.fits")

    dark = ma.masked_outside(fits.getdata(dfile), MIN_DARK, MAX_DARK)
    flat = ma.masked_outside(fits.getdata(ffile), MIN_FLAT, MAX_FLAT)

    pipe = RawPipeline(calibrate, {"dark":dark, "flat":flat})
    pipe.run()
```

# Data calibration -2

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```
if __name__ == "__main__":
    dfile = os.path.join(DATADIR, "masterdark.fits")
    ffile = os.path.join(DATADIR, "masterflat.fits")
```

  ↑ Direct execution

```
    dark = ma.masked_outside(fits.getdata(dfile), MIN_DARK, MAX_DARK)
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```

  ↑ Set up file names for *dark* e *flat*

```
    pipe = RawPipeline(calibrate, {"dark":dark, "flat":flat})
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# Data calibration -2

A complete example - 8

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if __name__ == "__main__":
    dfile = os.path.join(DATADIR, "masterdark.fits")
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dark = ma.masked_outside(fits.getdata(dfile), MIN_DARK, MAX_DARK)
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```

Set up file names for dark e flat

```
pipe = RawPipeline(calibrate, {"dark":dark, "flat":flat})
```

Read dark and flat images.

Function **masked\_outside()** converts the image into a *masked array*, masking out “unreasonable” pixels (i.e.: pixels with value outside the specified interval).

```
pipe.run()
```

# Data calibration -2

A complete example - 8

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if __name__ == "__main__":
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dark = ma.masked_outside(fits.getdata(dfile), MIN_DARK, MAX_DARK)
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Set up file names for *dark* e *flat*

Read *dark* and *flat* images.

Function **masked\_outside()** converts the image into a *masked array*, masking out “unreasonable” pixels (i.e.: pixels with value outside the specified interval).

```
pipe = RawPipeline(calibrate, {"dark":dark,
                                "flat":flat})
pipe.run()
```

Create an object *RawPipeline*

# Data calibration -2

A complete example - 8

Let's start a description of the procedure from the bottom of file calibrate.py

```
if __name__ == "__main__":
    dfile = os.path.join(DATADIR, "masterdark.fits")
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Read *dark* and *flat* images.

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pipe = RawPipeline(calibrate, {"dark":dark, "flat":flat})
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Create an object *RawPipeline*

Fire the loop on raw data files

# Data calibration -2

A complete example - 8

Let's start a description of the procedure from the bottom of file calibrate.py

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if __name__ == "__main__":
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Create an object *RawPipeline*

Fire the loop on raw data files

Being *dark* and *flat* masked arrays, any operations involving them will be performed ignoring non valid pixels.

# Data calibration -2

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if __name__ == "__main__":
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# Data calibration -3

A complete example - 9

Let's see the actual calibration function in file  
calibrate.py

```
def calibrate(img, step, args, common_args):
    shortname = os.path.basename(args["filename"])
    print("--- Calibrating:", shortname, end=" ", flush=True)

    cal = (img-common_args["dark"])/common_args["flat"]

    calname = os.path.join("work", shortname.replace("fit", "cal"))

    cal.dump(calname)
    print("OK")
```

# Data calibration -3

A complete example - 9

Let's see the actual calibration function in file `calibrate.py`

```
def calibrate(img, step, args, common_args):
    shortname = os.path.basename(args["filename"])
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```

Some *output* to show that  
the procedure is running

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def calibrate(img, step, args, common_args):  
    shortname = os.path.basename(args["filename"])  
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Some *output* to show that the procedure is running

```
    cal = (img-common_args["dark"])/common_args["flat"]
```

Computing calibrated image. Expression's arguments are *arrays*, thus the operations are performed element by element.

```
    calname = os.path.join("work", shortname.replace("fit", "cal"))
```

```
    cal.dump(calname)  
    print("OK")
```

# Data calibration -3

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    calname = os.path.join("work", shortname.replace("fit", "cal"))

    cal.dump(calname)
    print("OK")
```

Some output to show that the procedure is running

Computing calibrated image. Expression's arguments are arrays thus the operations are performed element by element.   
 dark and flat are *masked arrays*, thus non valid element are not considered and the result is again a *masked array*.

# Data calibration -3

A complete example - 9

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def calibrate(img, step, args, common_args):
    shortname = os.path.basename(args["filename"])
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    cal.dump(calname)
    print("OK")
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Some output to show that the procedure is running

Computing calibrated image. Expression's arguments are arrays thus the operations are element-wise. ~~dark and flat are masked arrays, thus non valid element are not considered and the result is again a masked array.~~

dark and flat are *masked arrays*, thus non valid element are not considered and the result is again a *masked array*.

Generate a proper file name for the calibrated image

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A complete example - 9

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Generate a proper file name for the calibrated image

Write calibrated image onto file

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A complete example - 9

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```

Some output to show that the procedure is running

Computing calibrated image. Expression's arguments are arrays thus the operations are performed element by element. Since dark and flat are *masked arrays*, thus non valid elements are not considered and the result is again a *masked array*.

Generate a proper file name for the calibrated image

Write calibrated image onto file

Reminder: function **calibrate()** is called once for each data file in the list.

# Data calibration -3

A complete example - 9

Let's see the actual calibration function in file `calibrate.py`

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def calibrate(img, step, args, common_args):
    shortname = os.path.basename(args["filename"])
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Generate a proper file name for the calibrated image

Write calibrated image onto file

Reminder: function **calibrate()** is called once for each data file in the list.

## Running `calibrate.py`

```
$ python calibrate.py -d /dati-18.7.2017 -g img-\*
```

### file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
        self.ax = self.fig.gca()
        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

    if __name__ == "__main__":
        img = np.load(sys.argv[1])
        sel = ObjectSelector(img)
        sel.start()
        with open("selected.dat", "w") as fpt:
            print(sel.circles, file=fpt)
```

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
        self.ax = self.fig.gca()
        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

    if __name__ == "__main__":
        img = np.load(sys.argv[1])
        sel = ObjectSelector(img)
        sel.start()
        with open("selected.dat", "w") as fpt:
            print(sel.circles, file=fpt)
```

Find candidate stars

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()           ← Set interactive mode
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
        self.ax = self.fig.gca()
        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, img):
        plt.ion()
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        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
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            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
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            self.circles.append([self.xc, self.yc, RADIUS])
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if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
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        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to corresponding functions

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
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        show(image)
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        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to execution loop.

Execution loop.  
Note: `plt.pause()`

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
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        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to execution loop.

Execution loop.  
Note: `plt.pause()`

Function called on button press

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, img):
        plt.ion()
        self.fig = plt.figure()
        show(img)
        plt.title("Left click to select object. End with right click")
        self.ax = self.fig.gca()
        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
        cid2 = self.fig.canvas.mpl_connect("button_release_event", self._release)
        self.goon = True
        self.circles = []
        plt.show()

    def start(self):
        while self.goon: plt.pause(0.10)

    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
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```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to execution loop.

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, img):
        plt.ion()
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        show(img)
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        elif event.button == 3:
            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to execution loop.

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

- Generate a circle  
- Save position  
- Draw it

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
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        self.ax = self.fig.gca()
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        self.goon = True
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    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
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if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to execution loop.

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

- Generate a circle  
- Save position  
- Draw it

Program execution

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
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        self.goon = True
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    def start(self):
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    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

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            self.goon = False

if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to execution loop.

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

- Generate a circle  
- Save position  
- Draw it

Get image name from command line

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
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        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
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        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

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        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
            self.ax.add_artist(circle)
            self.circles.append([self.xc, self.yc, RADIUS])
            plt.draw()
        elif event.button == 3:
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if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to functions

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

- Generate a circle  
- Save position  
- Draw it

Get image name from command line

Instantiate ObjectSelector

# Stars selection -1

A complete example - 10

file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
        self.ax = self.fig.gca()
        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
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        elif event.button == 3:
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if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to functions

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

- Generate a circle  
- Save position  
- Draw it

Get image name from command line

Start the display loop

# Stars selection -1

A complete example - 10

## file: select.py

```
import sys
import numpy as np
import matplotlib.pyplot as plt
from show import show

RADIUS = 50
class ObjectSelector:
    def __init__(self, image):
        plt.ion()
        self.fig = plt.figure()
        show(image)
        plt.title("Left click to select object. End with right click")
        self.ax = self.fig.gca()
        cid1 = self.fig.canvas.mpl_connect("button_press_event", self._press)
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    def start(self):
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    def _press(self, event):
        if event.button == 1: self.xc = event.xdata; self.yc = event.ydata

    def _release(self, event):
        if event.button == 1:
            circle = plt.Circle((self.xc, self.yc), RADIUS, color="cyan", fill=False)
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if __name__ == "__main__":
    img = np.load(sys.argv[1])
    sel = ObjectSelector(img)
    sel.start()
    with open("selected.dat", "w") as fpt:
        print(sel.circles, file=fpt)
```

Find candidate stars

Set interactive mode

Show the image for star selection

Connect events to functions

Execution loop.  
Note: `plt.pause()`

Function called on button press

Function called on button release

- Generate a circle  
- Save position  
- Draw it

Get image name from command line

Start the display loop

Eventually, write circle positions into a file

# Stars selection -2

A complete example - 11

How does `select.py` work

\$: `python select.py work/img-460.cal`

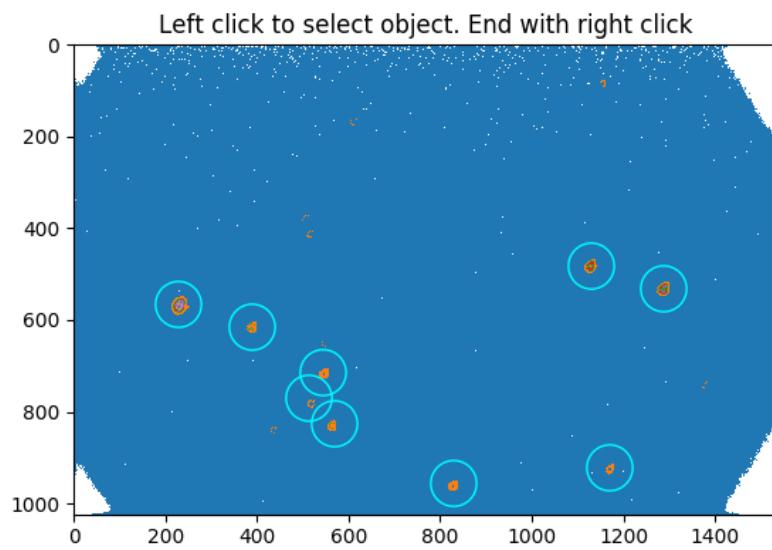
# Stars selection -2

A complete example - 11

## How does select.py work

```
$: python select.py work/img-460.cal
```

- Left Click on stars to add to selection



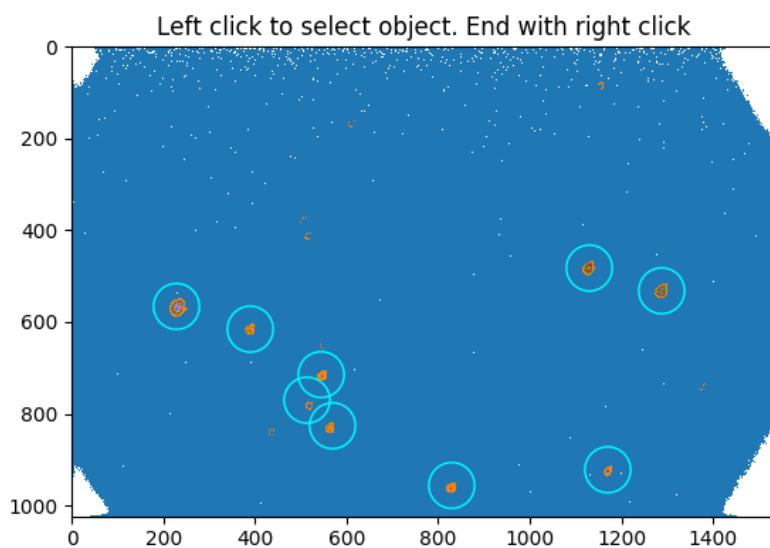
# Stars selection -2

A complete example - 11

## How does select.py work

```
$: python select.py work/img-460.cal
```

- Left *Click* on stars to add to selection



- To terminate: right *click*

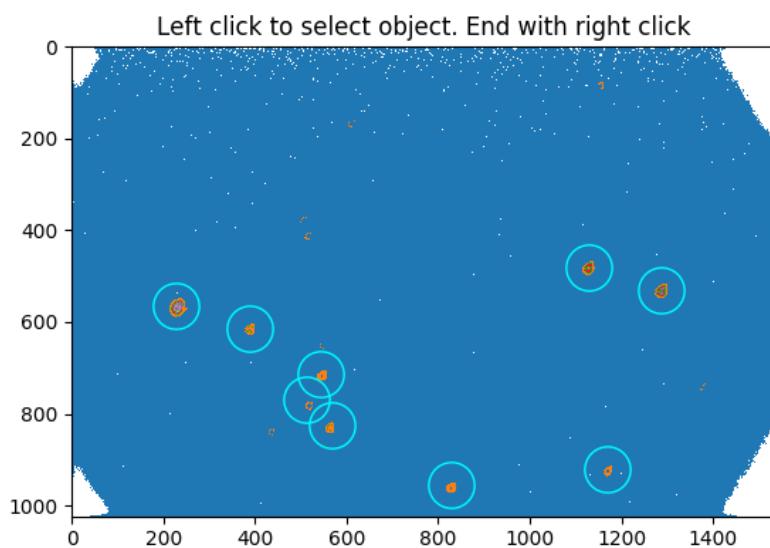
# Stars selection -2

A complete example - 11

## How does select.py work

\$: `python select.py work/img-460.cal`

- Left *Click* on stars to add to selection



- To terminate: right *click*
- Positions are written into file `selected.dat` as:

---

```
[[228.66129032258061, 566.00322580645161, 50], [1129.8225806451612, 482.39032258 064515, 50], [1287.758064516129, 531.9387096774193, 50], [389.69354838709671, 615.55161290322576, 50], [544.5322580645161, 714.64838709677417, 50], [513.5645161290322, 770.39032258064515, 50], [569.30645161290317, 826.13225806451612, 50], [ 829.4354838709678, 956.19677419354844, 50], [1170.0806451612902, 922.13225806451612, 50]]
```

---

Defining position of selected circular areas (nine, in the example)

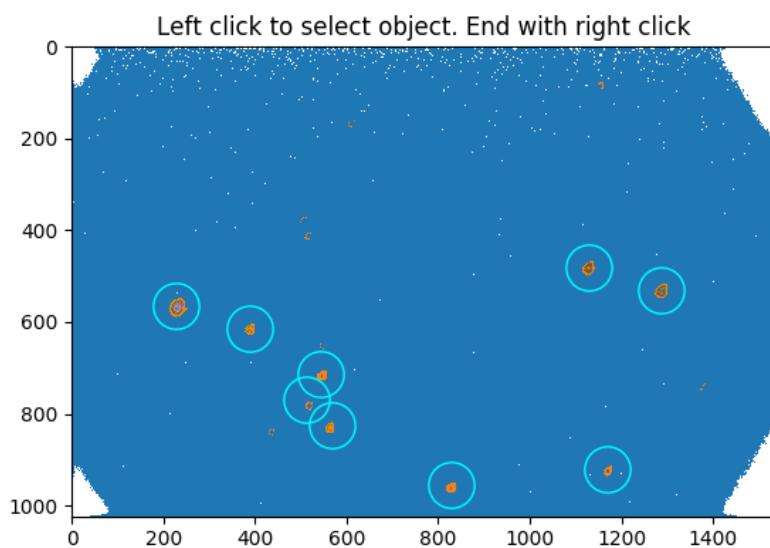
# Stars selection -2

A complete example - 11

## How does select.py work

\$: `python select.py work/img-460.cal`

- Left *Click* on stars to add to selection



- To terminate: right *click*
- Positions are written into file `selected.dat` as:

---

```
[[228.66129032258061, 566.00322580645161, 50], [1129.8225806451612, 482.39032258 064515, 50], [1287.758064516129, 531.9387096774193, 50], [389.69354838709671, 615.55161290322576, 50], [544.5322580645161, 714.64838709677417, 50], [513.5645161290322, 770.39032258064515, 50], [569.30645161290317, 826.13225806451612, 50], [ 829.4354838709678, 956.19677419354844, 50], [1170.0806451612902, 922.13225806451612, 50]]
```

---

Defining position of selected circular areas (nine, in the example)



Program select.py was maintained as simple as possible. In “real life” some enhancements would be necessary:

- Make it clear that the exoplanet’s star must be selected as first.

Program select.py was maintained as simple as possible. In “real life” some enhancements would be necessary:

- Make it clear that the exoplanet’s star must be selected as first.
- Make cross size proportional to star brightness

Program select.py was maintained as simple as possible. In “real life” some enhancements would be necessary:

- Make it clear that the exoplanet’s star must be selected as first.
- Make cross size proportional to star brightness
- Make error messages more explanatory

Program select.py was maintained as simple as possible. In “real life” some enhancements would be necessary:

- Make it clear that the exoplanet’s star must be selected as first.
- Make cross size proportional to star brightness
- Make error messages more explanatory
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The following simple procedure gives an idea of the centering problem.

Note: it uses again a Pipeline object.

### file: timelapse.py

```
import json, os
import matplotlib.pyplot as plt
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CIRCFILE = "selected.dat"
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    axes = common_args["axes"]
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    show(img)
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if __name__ == "__main__":
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    fig = plt.figure()
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Function called on each data file



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### file: timelapse.py

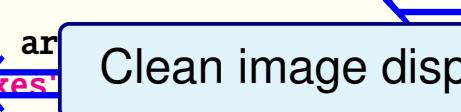
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```

Function called on each data file

Clean image display



# Stars centering -1

A complete example - 13

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Function called on each data file

Show the new image

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A complete example - 13

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Function called on each data file

Show the new image

Draw circles

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Add title

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Show the new image

Draw circles

Add title

Visualization pause

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Draw circles

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Visualization pause

Set matplotlib interactive mode

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Function called on each data file

Show the new image

Draw circles

Add title

Visualization pause

Set matplotlib interactive mode

Generate circles based on data from file

# Stars centering -1

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Visualization pause

Set matplotlib interactive mode

Generate circles based on data from file

Create the CalPipeline object

# Stars centering -1

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Function called on each data file

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Draw circles

Add title

Visualization pause

Set matplotlib interactive mode

Generate circles based on data from file

Create the CalPipeline object

Start loop

# Stars centering -1

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Show the new image

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Visualization pause

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Create the CalPipeline object

Start loop



### Resuming:

- Each image must be re-centered so that it covers the same area on sky.

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... in general:

- It could be necessary to take field rotation into account

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... in general:
- It could be necessary to take field rotation into account

# Doing photometry -1

A complete example - 15

file: photometry1.py

```
import json
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]

DIFF_FLUX = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)

if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
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# Doing photometry -1

A complete example - 15

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Import required tools, aperture_photometry

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```

Import required, aperture\_photometry  
Selected star positions are here.  
The first one is our target

# Doing photometry -1

A complete example - 15

file: photometry1.py

```
import json
from photutils import centroid_ellipse
from pipeline import CalPipeline
CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
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Import required, `aperture_photometry`

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Function `flux()` computes  
the total flux in a circular  
area around the given point  
`cc`

# Doing photometry -1

A complete example - 15

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Function `calc_flux()` computes  
the differential flux of  
target star for each input  
image

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    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]
```

Function `flux()` computes  
the total flux in a circular  
area around the given point  
`cc`

```
DIFF_FLUX = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)
```

Function `calc_flux()` computes  
the differential flux of  
target star for each input  
image

```
if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

We use the Pipeline  
class to iterate on the  
calibrated images

# Doing photometry -1

A complete example - 15

file: photometry1.py

```
import json
from photutils import centroid_photometry
from pipeline import CalPipe
CIRCFILE = "selected.dat"
STAR_RADIUS = 30
```

Import required, `centroid_photometry`  
Selected star positions are here.  
The first one is our target

```
def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]
```

Function `flux()` computes  
the total flux in a circular  
area around the given point  
`cc`

```
DIFF_FLUX = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)
```

Function `calc_flux()` computes  
the differential flux of  
target star for each input  
image

```
if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

We use the Pipeline  
class to iterate on the  
calibrated images

**Note** Function `flux`, at each iteration, updates the position of the `CircularAperture()` object for the following step (because `aperture_photometry()` returns the centroid position together with the computed flux).

# Doing photometry -1

A complete example - 15

file: photometry1.py

```
import json
from photutils import centroid_photometry
from pipeline import CalPipe
CIRCFILE = "selected.dat"
STAR_RADIUS = 30
```

Import required, `aperture_photometry`

Selected star positions are here.  
The first one is our target

```
def flux(img, cc):
    aperture = CircularAperture(cc[0:2], radius=STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]
```

Function `flux()` computes  
the total flux in a circular  
area around the given point  
`cc`

```
DIFF_FLUX = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)
```

Function `calc_flux()` computes  
the differential flux of  
target star for each input  
image

```
if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

We use the Pipeline  
class to iterate on the  
calibrated images

**Note** Function `flux`, at each iteration, updates the position of the `CircularAperture()` object for the following step (because `aperture_photometry()` returns the centroid position together with the computed flux).



# Doing photometry -2

A complete example - 16

## Running photometry.py from ipython

---

```
In [1]: %run photometry.py -d work
```

```
Step 0 done
```

```
Step 1 done
```

```
....
```

```
Step 771 done
```

```
Step 772 done
```

```
In [2]: plt.plot(DIFF_FLUX)
```

---

# Doing photometry -2

A complete example - 16

## Running photometry.py from ipython

```
In [1]: %run photometry.py -d work
```

```
Step 0 done
```

```
Step 1 done
```

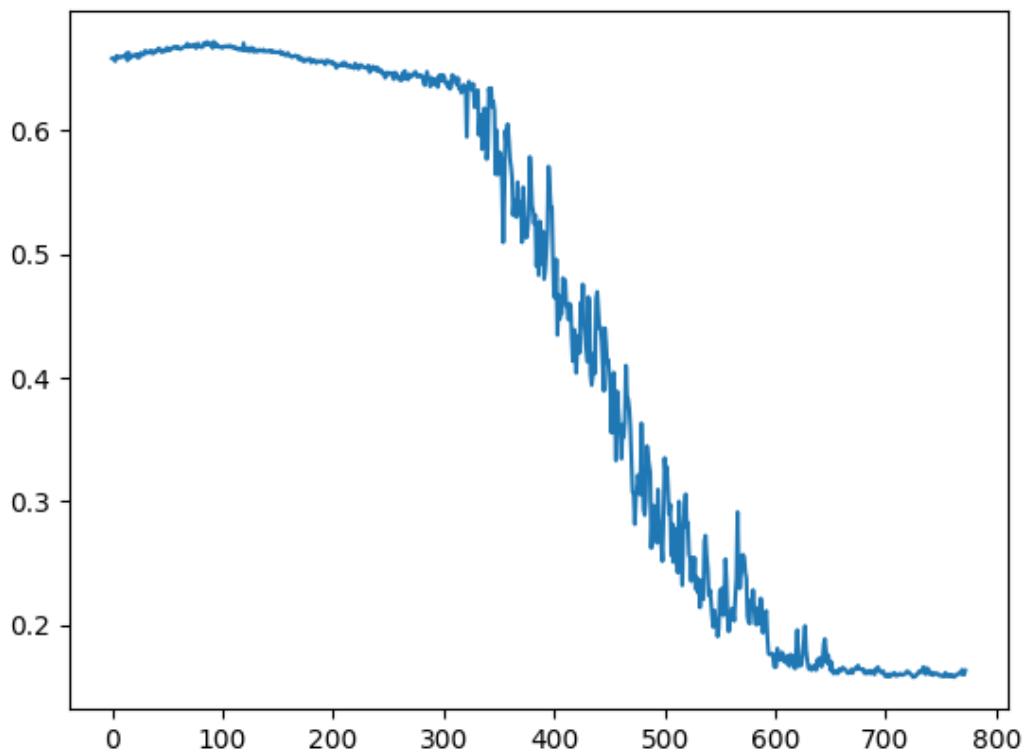
```
....
```

```
Step 771 done
```

```
Step 772 done
```

```
In [2]: plt.plot(DIFF_FLUX)
```

... and here's the result:



Clearly that's not what we would expect!

# Doing photometry -2

A complete example - 16

## Running photometry.py from ipython

```
In [1]: %run photometry.py -d work
```

```
Step 0 done
```

```
Step 1 done
```

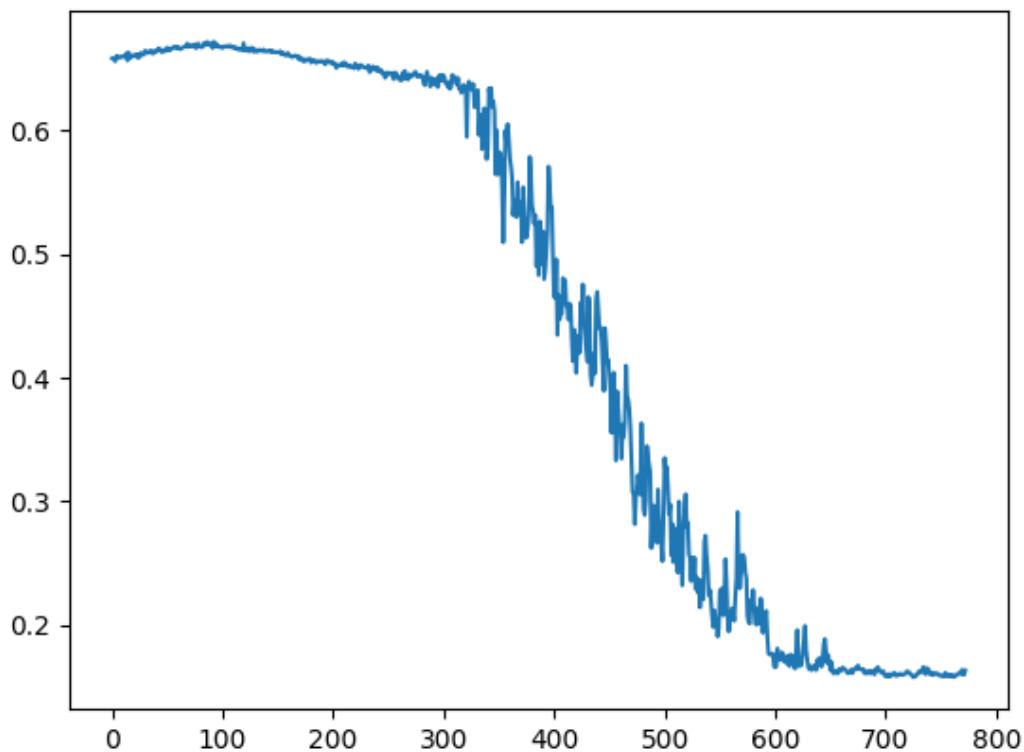
```
....
```

```
Step 771 done
```

```
Step 772 done
```

```
In [2]: plt.plot(DIFF_FLUX)
```

... and here's the result:



Clearly that's not what we would expect!



# Doing photometry -3

A complete example - 17

Let's look at single stars.

# Doing photometry -3

A complete example - 17

Let's look at single stars.

file: photometry1.py

```
import json
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]

DIFF_FLUX = []
STAR_FLUXES=[]
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    STAR_FLUXES.append(fluxes)
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)

if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

# Doing photometry -3

A complete example - 17

Let's look at single stars.

file: photometry1.py

```
import json
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]

DIFF_FLUX = []
STAR_FLUXES = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    STAR_FLUXES.append(fluxes)
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)

if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

photometry1.py is very similar to the previous function

# Doing photometry -3

A complete example - 17

Let's look at single stars.

file: photometry1.py

```
import json
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]

DIFF_FLUX = []
STAR_FLUXES = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    STAR_FLUXES.append(fluxes) ← but it also stores fluxes
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)

if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

photometry1.py is very similar to the previous function

but it also stores fluxes for each star

# Doing photometry -3

A complete example - 17

Let's look at single stars.

file: photometry1.py

```
import json
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = flx["xcenter"][0].value
    cc[1] = flx["ycenter"][0].value
    return flx["aperture_sum"][0]

DIFF_FLUX = []
STAR_FLUXES = []
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    STAR_FLUXES.append(fluxes) ← but it also stores fluxes
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)

if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

photometry1.py is very similar to the previous function

but it also stores fluxes for each star



# Doing photometry -4

A complete example - 18

---

```
In [30]: %run photometry1.py -d work
```

```
In [31]: stars = np.array(STAR_FLUXES).transpose()
```

```
In [32]: plt.plot(stars[0])
```

```
In [33]: plt.plot(stars[1])
```

```
In [34]: plt.plot(stars[2])
```

```
In [35]: plt.plot(stars[3])
```

---

# Doing photometry -4

A complete example - 18

---

```
In [30]: %run photometry1.py -d work
```

```
In [31]: stars = np.array(STAR_FLUXES).transpose()
```

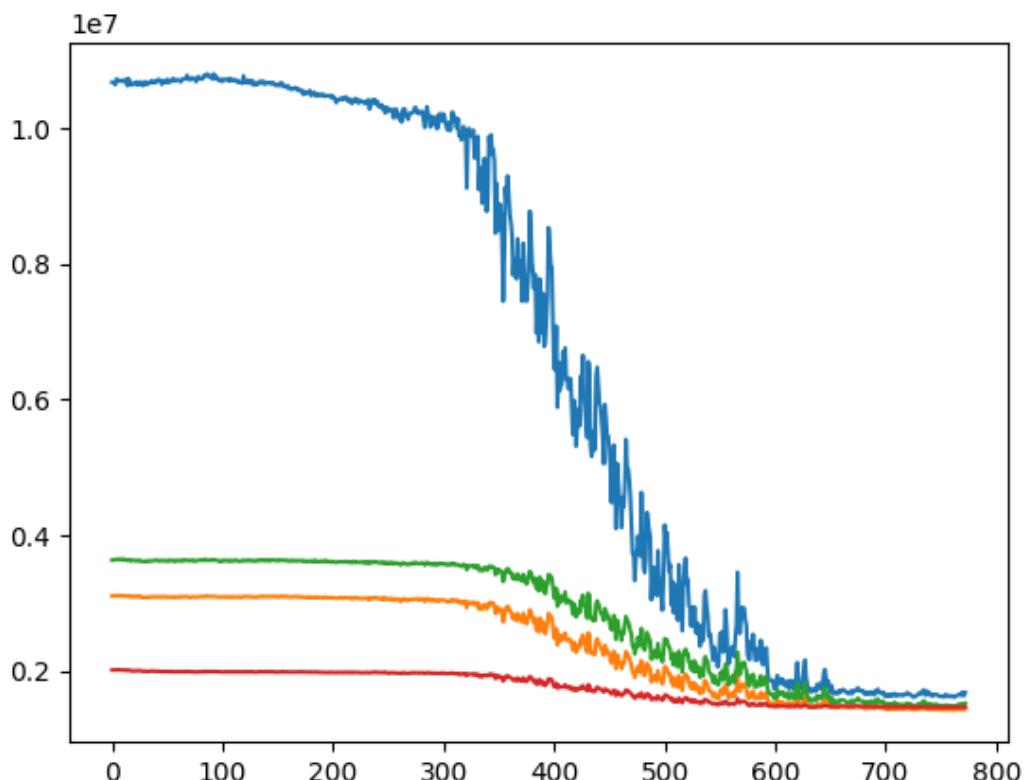
```
In [32]: plt.plot(stars[0])
```

```
In [33]: plt.plot(stars[1])
```

```
In [34]: plt.plot(stars[2])
```

```
In [35]: plt.plot(stars[3])
```

---



My first interpretation of the above plot: the recentering algorithm is not working.

I.e.: starting around the 300<sup>th</sup> image the stars drop out of the aperture

# Doing photometry -5

A complete example - 19

Let's see it better:

file: timelapse1.py

---

```
import json, os
import matplotlib.pyplot as plt
from show import show
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30
def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    ccx = flx["xcenter"][0].value
    ccy = flx["ycenter"][0].value
    plt.title("img: %d - Center: %.2f, %.2f / %.2f" %(step, cc[0], cc[1], cc[2]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)

if __name__ == "__main__":
    plt.ion()
    fig = plt.figure()
    ax = fig.gca()
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(time_lapse, {"axes":ax, "center":cdata[0]})
    pipe.run()
```

---

# Doing photometry -5

A complete example - 19

Let's see it better:

file: timelapse1.py

```
import json, os
import matplotlib.pyplot as plt
from show import show
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30

def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    ccx = flx["xcenter"][0].value
    ccy = flx["ycenter"][0].value
    plt.title("img: %d - Center: %.2f, %.2f / %.2f" %(step, cc[0], cc[1], cc[2]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)

if __name__ == "__main__":
    plt.ion()
    fig = plt.figure()
    ax = fig.gca()
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(time_lapse, {"axes":ax, "center":cdata[0]})
    pipe.run()
```

timelapse1.py merges code from timelapse.py and photometry.py to plot the centroid position at each step

# Doing photometry -5

A complete example - 19

Let's see it better:

file: timelapse1.py

```
import json, os
import matplotlib.pyplot as plt
from show import show
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30

def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    ccx = flx["xcenter"][0].value
    ccy = flx["ycenter"][0].value
    plt.title("img: %d - Center: %.2f, %.2f / %.2f" %(step, cc[0], cc[1], cc[2]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)

if __name__ == "__main__":
    plt.ion()
    fig = plt.figure()
    ax = fig.gca()
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(time_lapse, {"axes":ax, "center":cdata[0]})
    pipe.run()
```

timelapse1.py merges code from timelapse.py and photometry.py to plot the centroid position at each step

Notably it shows the aperture center and the computed centroid

# Doing photometry -5

A complete example - 19

Let's see it better:

file: timelapse1.py

```
import json, os
import matplotlib.pyplot as plt
from show import show
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30

def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    ccx = flx["xcenter"][0].value
    ccy = flx["ycenter"][0].value
    plt.title("img: %d - Center: %.2f, %.2f / %.2f, %.2f" %(step, cc[0], cc[1], ccx, ccy))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)

if __name__ == "__main__":
    plt.ion()
    fig = plt.figure()
    ax = fig.gca()
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(time_lapse, {"axes":ax, "center":cdata[0]})
    pipe.run()
```

timelapse1.py merges code from timelapse.py and photometry.py to plot the centroid position at each step

Notably it shows the aperture center and the computed centroid

---

```
python timelapse1.py -d work
```

---

# Doing photometry -5

A complete example - 19

Let's see it better:

file: timelapse1.py

```
import json, os
import matplotlib.pyplot as plt
from show import show
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30

def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    ccx = flx["xcenter"][0].value
    ccy = flx["ycenter"][0].value
    plt.title("img: %d - Center: %.2f, %.2f / %.2f" %(step, cc[0], cc[1], cc[2]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)

if __name__ == "__main__":
    plt.ion()
    fig = plt.figure()
    ax = fig.gca()
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(time_lapse, {"axes":ax, "center":cdata[0]})
    pipe.run()
```

timelapse1.py merges code from timelapse.py and photometry.py to plot the centroid position at each step

Notably it shows the aperture center and the computed centroid

python

In conclusion: RTFM (Read the f...  
manual)

# Doing photometry -5

A complete example - 19

Let's see it better:

file: timelapse1.py

```
import json, os
import matplotlib.pyplot as plt
from show import show
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30

def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    aperture = CircularAperture(cc[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    ccx = flx["xcenter"][0].value
    ccy = flx["ycenter"][0].value
    plt.title("img: %d - Center: %.2f, %.2f / %.2f" %(step, cc[0], cc[1], cc[2]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)

if __name__ == "__main__":
    plt.ion()
    fig = plt.figure()
    ax = fig.gca()
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(time_lapse, {"axes":ax, "center":cdata[0]})
    pipe.run()
```

timelapse1.py merges code from timelapse.py and photometry.py to plot the centroid position at each step

Notably it shows the aperture center and the computed centroid

python

In conclusion: RTFM (Read the f...  
manual)



# Doing photometry -6

A complete example - 20

## Let's compute the centroid

file: centroid.py

```
from photutils import centroid_com

def centroid(img, ctr):
    x0 = int(max(ctr[0]-ctr[2], 0))
    x1 = int(min(ctr[0]+ctr[2], img.shape[1]-1))
    y0 = int(max(ctr[1]-ctr[2], 0))
    y1 = int(min(ctr[1]+ctr[2], img.shape[0]-1))
    subimg = img[y0:y1, x0:x1]
    ctd = centroid_com(subimg)
    return (x0+ctd[0], y0+ctd[1])
```

... and use it again in timelapse

file: timelapse2.py

```
...
from centroid import centroid

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30
def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    ccx, ccy = centroid(img, cc)
    plt.title("img: %d - Center: %.2f, %.2f / %.2f, %.2f"%(step, cc[0], cc[1], cc[0], cc[1]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)
...
```

# Doing photometry -6

A complete example - 20

Let's compute the centroid

file: centroid.py

---

```
from photutils import centroid_com

def centroid(img, ctr):
    x0 = int(max(ctr[0]-ctr[2], 0))
    x1 = int(min(ctr[0]+ctr[2], img.shape[1]-1))
    y0 = int(max(ctr[1]-ctr[2], 0))
    y1 = int(min(ctr[1]+ctr[2], img.shape[0]-1))
    subimg = img[y0:y1, x0:x1]
    ctd = centroid_com(subimg)
    return (x0+ctd[0], y0+ctd[0])
```

---

... and use it again in timelapse

file: timelapse2.py

---

```
...
from centroid import centroid

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30
def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    ccx, ccy = centroid(img, cc)
    plt.title("img: %d - Center: %.2f, %.2f / %.2f, %.2f"%(step, cc[0], cc[1], cc[0], cc[1]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)
...
```

---

```
python timelapse2.py -d work
```

# Doing photometry -6

A complete example - 20

Let's compute the centroid

file: centroid.py

---

```
from photutils import centroid_com

def centroid(img, ctr):
    x0 = int(max(ctr[0]-ctr[2], 0))
    x1 = int(min(ctr[0]+ctr[2], img.shape[1]-1))
    y0 = int(max(ctr[1]-ctr[2], 0))
    y1 = int(min(ctr[1]+ctr[2], img.shape[0]-1))
    subimg = img[y0:y1, x0:x1]
    ctd = centroid_com(subimg)
    return (x0+ctd[0], y0+ctd[0])
```

---

... and use it again in timelapse

file: timelapse2.py

---

```
...
from centroid import centroid

CIRCFILE = "selected.dat"
BRACE = 30
STAR_RADIUS = 30
def time_lapse(img, step, args, common_args):
    axes = common_args["axes"]
    axes.cla()
    show(img)
    cc = common_args["center"]
    plt.plot((cc[0]-BRACE, cc[0]+BRACE), (cc[1], cc[1]), color="cyan")
    plt.plot((cc[0], cc[0]), (cc[1]-BRACE, cc[1]+BRACE), color="cyan")
    ccx, ccy = centroid(img, cc)
    plt.title("img: %d - Center: %.2f, %.2f / %.2f, %.2f"%(step, cc[0], cc[1], cc[0], cc[1]))
    cc[0] = ccx
    cc[1] = ccy
    plt.pause(0.10)
...
```

---

```
python timelapse2.py -d work
```

---

# Doing photometry -7

A complete example - 21

Here's the new version:

file: photometry2.py

```
import json
from photutils import centroid_com, CircularAperture, aperture_photometry
from pipeline import CalPipeline

from centroid import centroid

CIRCFILE = "selected.dat"
STAR_RADIUS = 30

def flux(img, cc):
    ctd = centroid(img, cc)
    aperture = CircularAperture(ctd[0:2], STAR_RADIUS)
    flx = aperture_photometry(img, aperture)
    cc[0] = ctd[0]
    cc[1] = ctd[1]
    return flx["aperture_sum"][0]

DIFF_FLUX = []
STAR_FLUXES=[]
def calc_flux(img, step, args, common_args):
    global DIFF_FLUX
    fluxes = []
    for cc in common_args["circles"]:
        fluxes.append(flux(img, cc))
    STAR_FLUXES.append(fluxes)
    diff_flux = fluxes[0]/sum(fluxes[1:])
    DIFF_FLUX.append(diff_flux)
    print("Step %d done"%step)

if __name__ == "__main__":
    with open(CIRCFILE) as cfile:
        cdata = json.load(cfile)
    pipe = CalPipeline(calc_flux, {"circles": cdata})
    pipe.run()
```

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We compute the centroid  
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Then we update the center position for the next step

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# Doing photometry -7

A complete example - 22

Let's try again:

# Doing photometry -7

A complete example - 22

Let's try again:

---

```
In [1]: %run photometry2.py -d work
```

```
Step 0 done
```

```
Step 1 done
```

```
...
```

```
Step 771 done
```

```
Step 772 done
```

---

```
In [2]: stars=np.array(STAR_FLUXES).transpose()
```

```
In [3]: plt.plot(stars[0])
```

```
In [4]: plt.plot(stars[1])
```

```
In [5]: plt.plot(stars[2])
```

---

```
In [6]: plt.plot(stars[3])
```

# Doing photometry -7

A complete example - 22

Let's try again:

In [1]: %run photometry2.py -d work

Step 0 done

Step 1 done

...

Step 771 done

Step 772 done

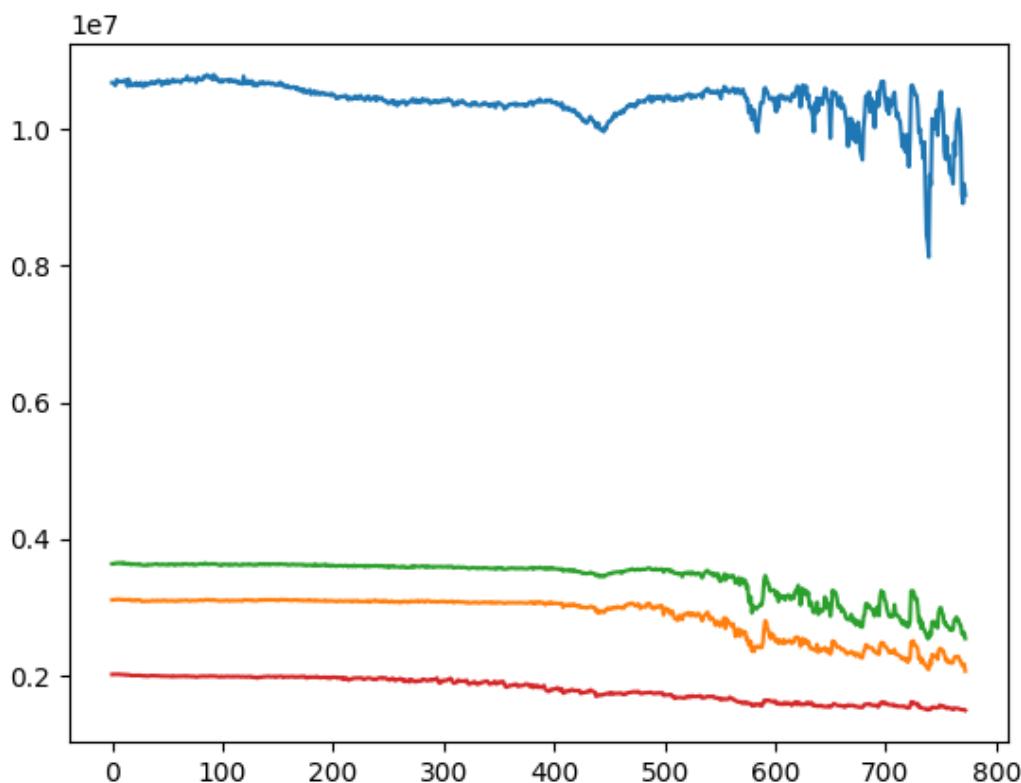
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# Doing photometry -8

A complete example - 23

And the new photometry curve:

# Doing photometry -8

A complete example - 23

And the new photometry curve:

---

```
In [7]: plt.plot(DIFF_FLUX)
```

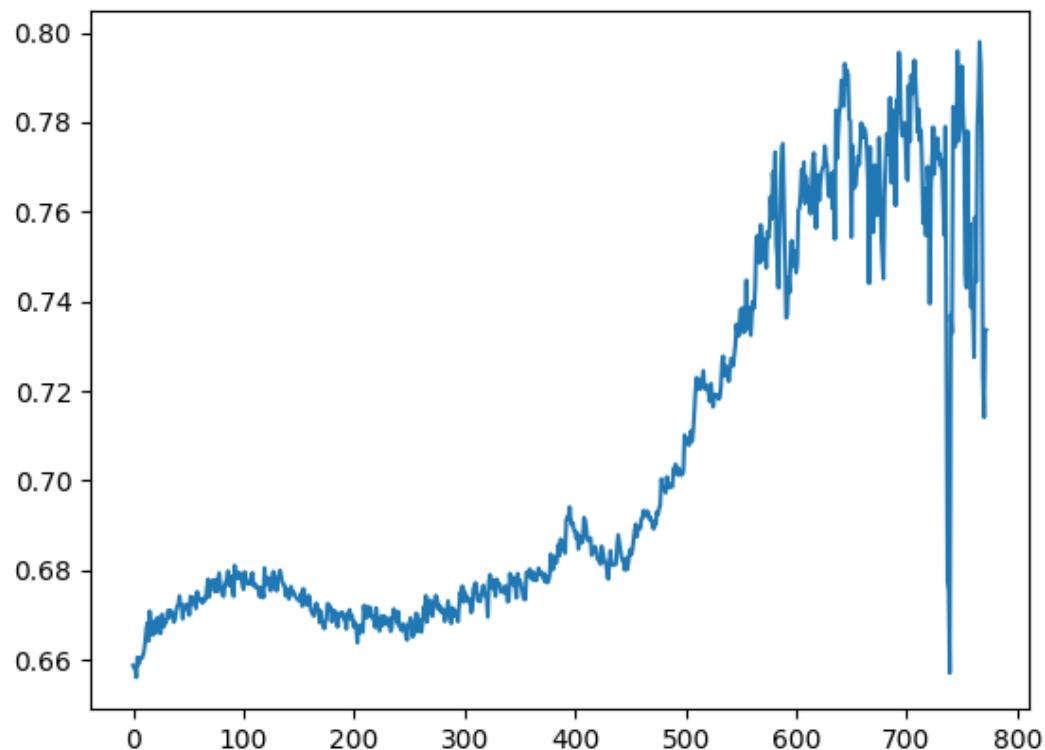
---

# Doing photometry -8

A complete example - 23

And the new photometry curve:

In [7]: `plt.plot(DIFF_FLUX)`

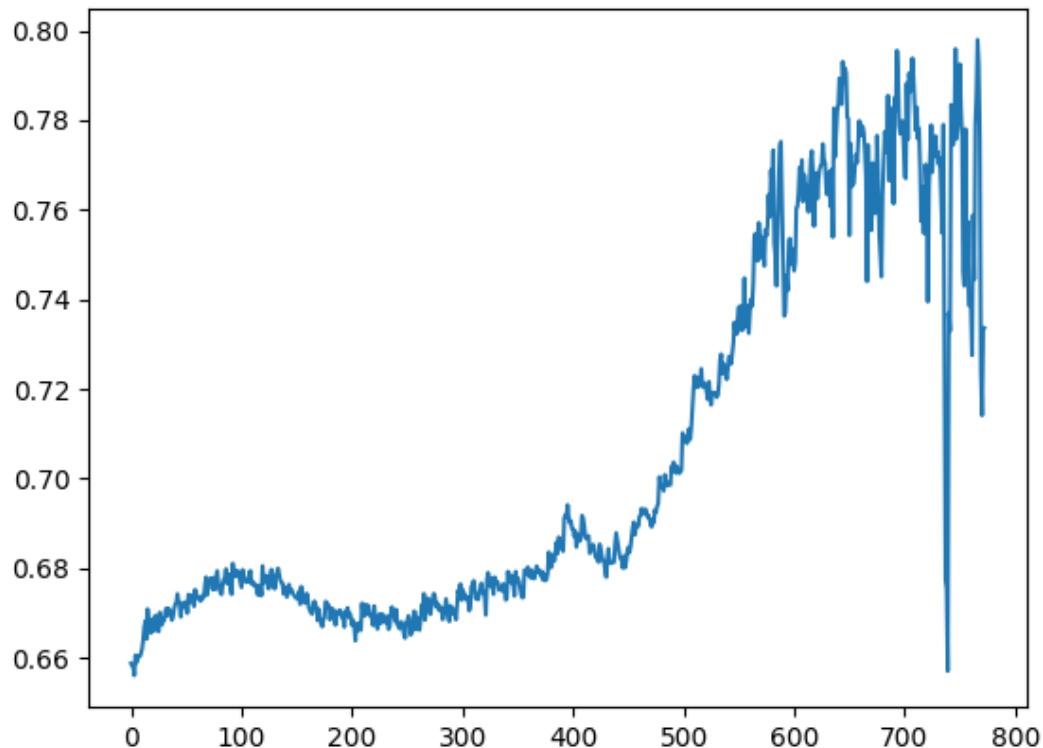


# Doing photometry -8

A complete example - 23

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In [7]: `plt.plot(DIFF_FLUX)`



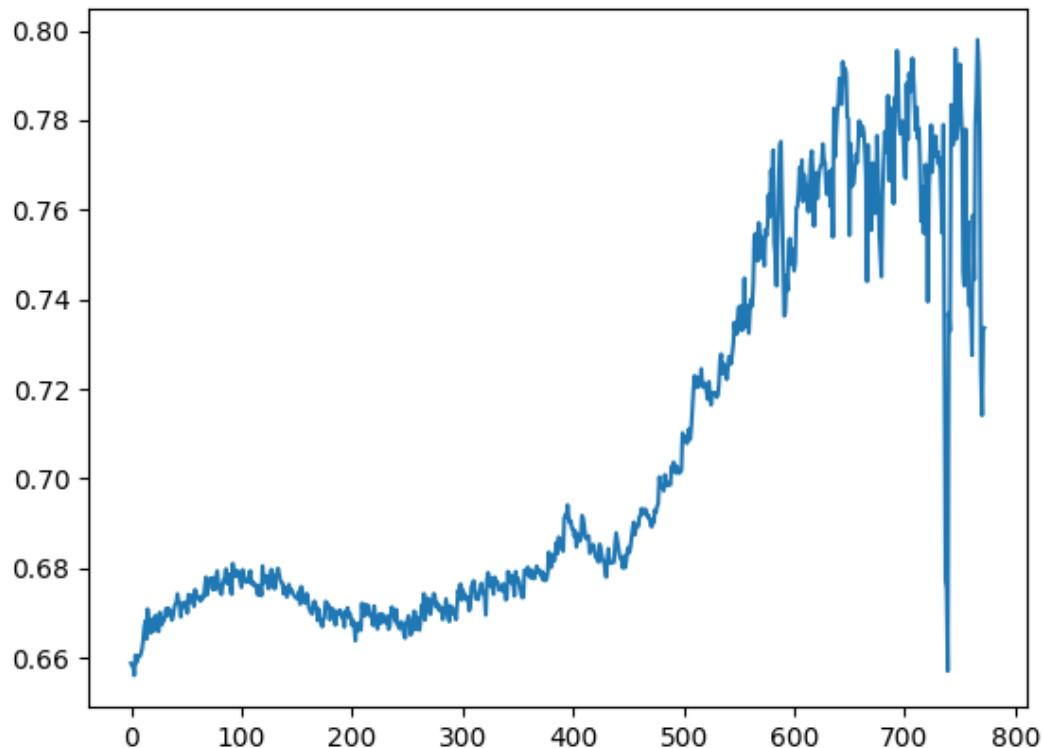
And clearly we're not there, yet ...

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A complete example - 23

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And clearly we're not there, yet ...



### What is next?

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## Another approach to recentering

## Another approach to recentering

I use the ObjectSelector defined earlier to find the best possible center of the target star in the first and last images.

```
In [1]: from myselect import ObjectSelector  
  
In [2]: img1=np.load("work/img-001.cal")  
  
In [3]: img2=np.load("work/img-773.cal")  
  
In [4]: sel = ObjectSelector(img1)  
  
In [5]: sel.start()  
  
In [6]: CIRC0=sel.circles[0]  
  
In [7]: sel = ObjectSelector(img2)  
  
In [8]: sel.start()  
  
In [9]: CIRC1=sel.circles[0]  
  
In [10]: SHX=np.linspace(CIRC0[0],CIRC1[0], 773)  
  
In [11]: SHY=np.linspace(CIRC0[1],CIRC1[1], 773)  
  
In [12]: SHX.dump("shx.dat")  
  
In [13]: SHY.dump("shy.dat")
```

### Another approach to recentering

I use the ObjectSelector defined earlier to find the best possible center first and last images.

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Load first and last images

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Generate a sequence of positions interpolating a straight line between the two points

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